

Week 9

Evolution of high-mass

November 16

Low-mass star lifecycle overview

Each step ★ has observable difference in L & T

What causes these changes?

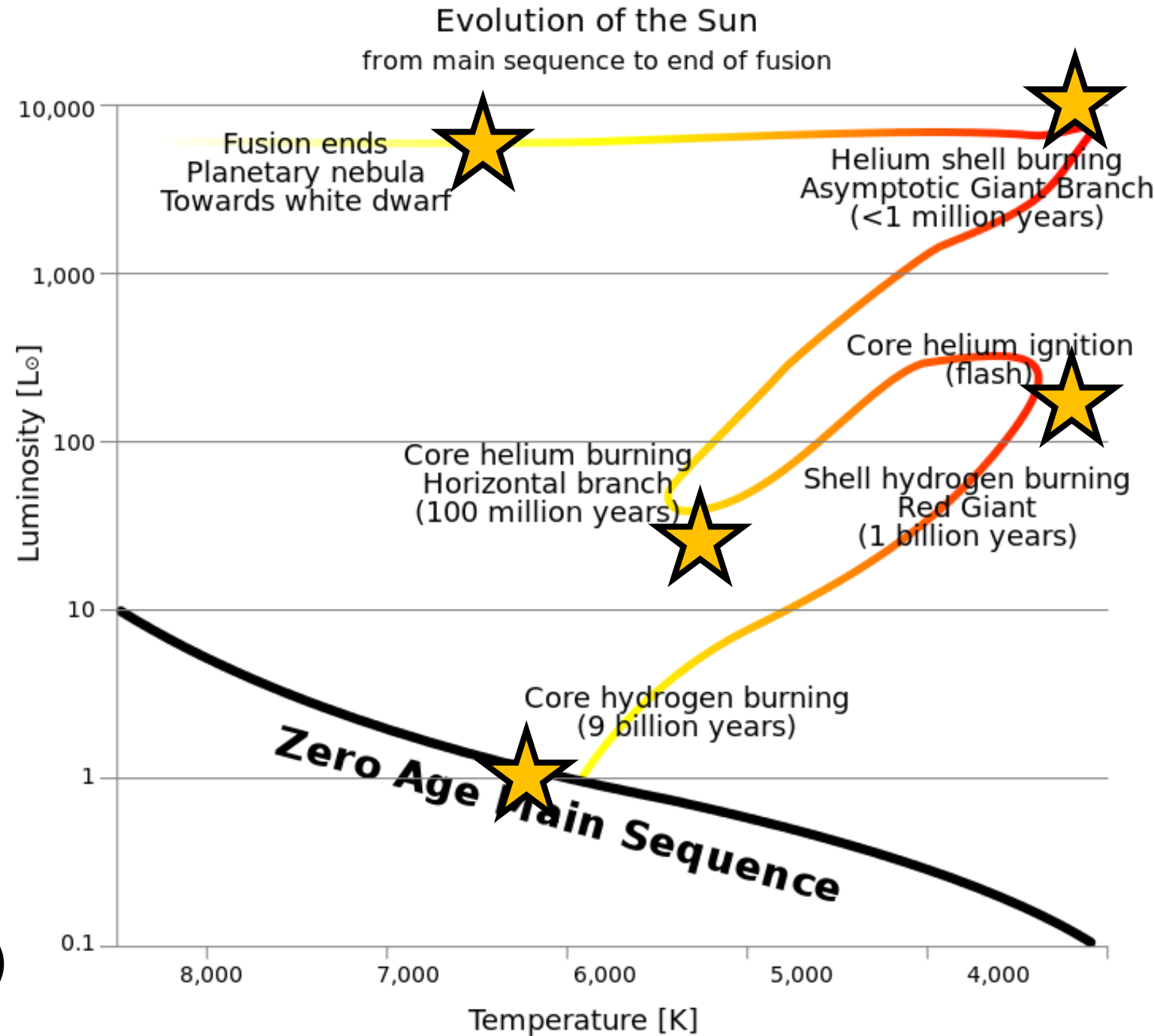
Changes in balance of:

- inward pull of gravity and
- outward pressure from core temperature

Why does balance change?

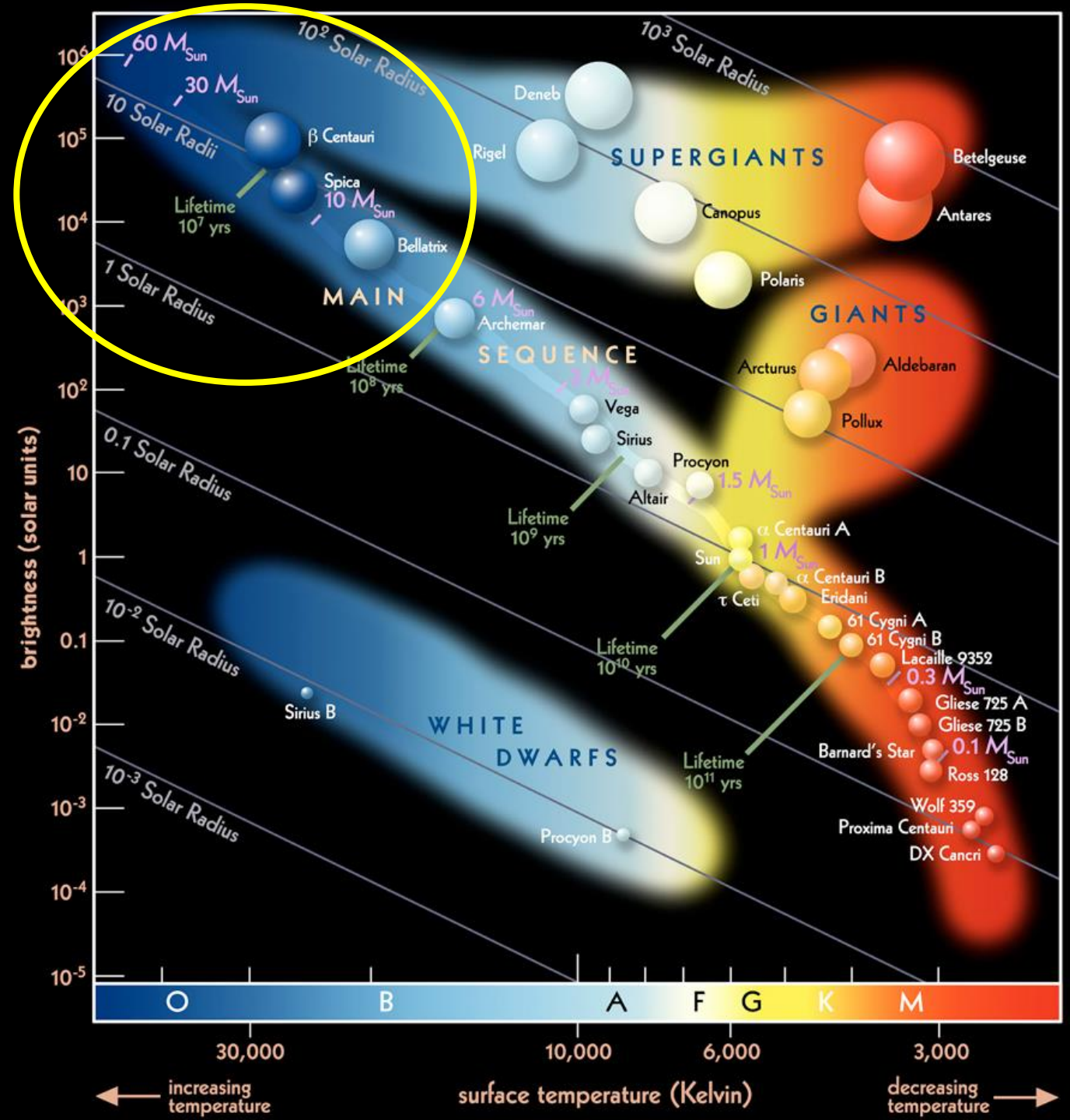
Changes in:

- Core composition/size
- Type of nuclear fuel (^1H or ^4He)

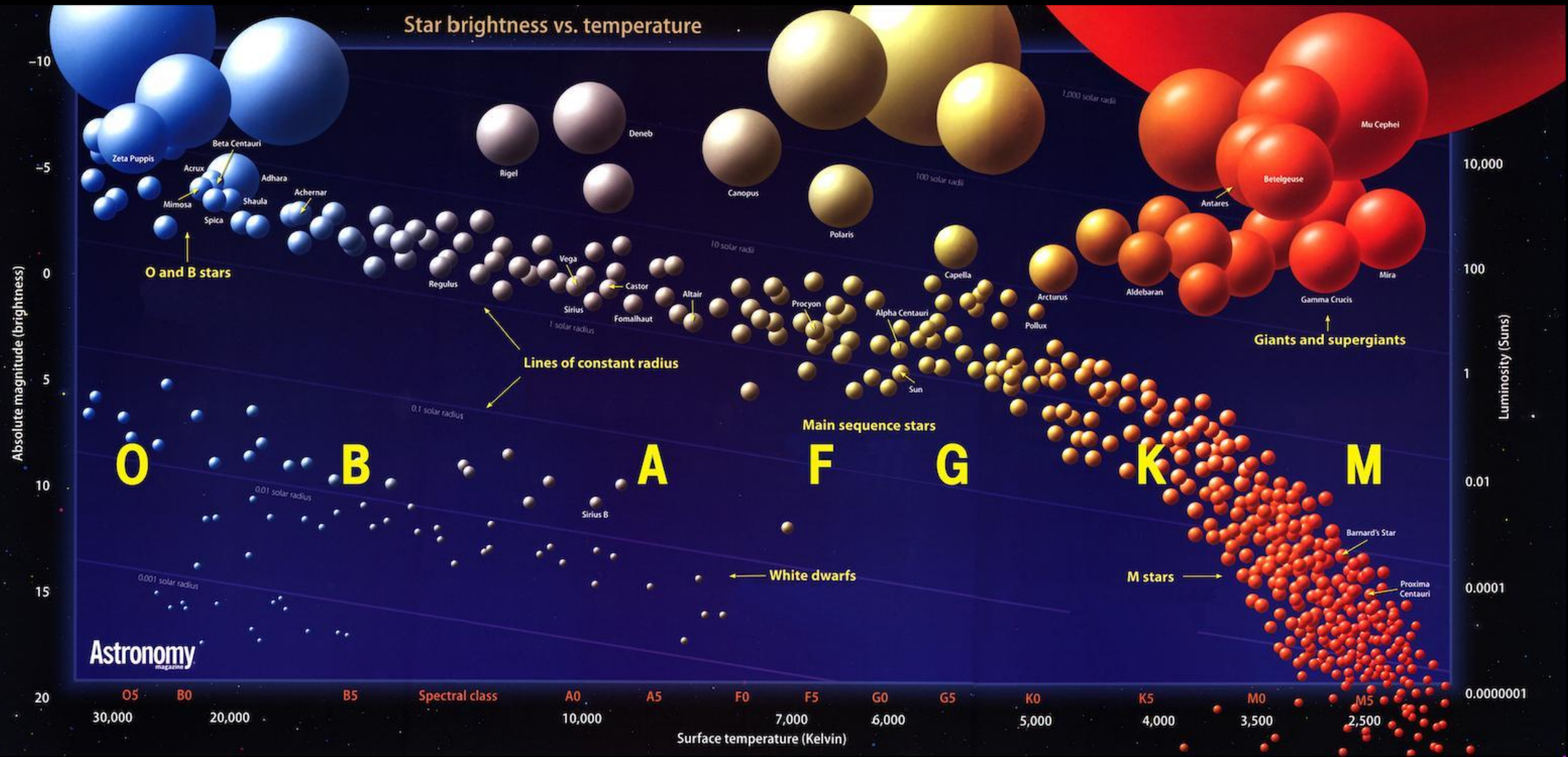


What about high-mass stars?

- Mass $> 8M_{\odot}$
- High T
- High L
- Large radius ($\sim 10 R_{\odot}$ while on main sequence)



What about high-mass stars?





Earth Moon



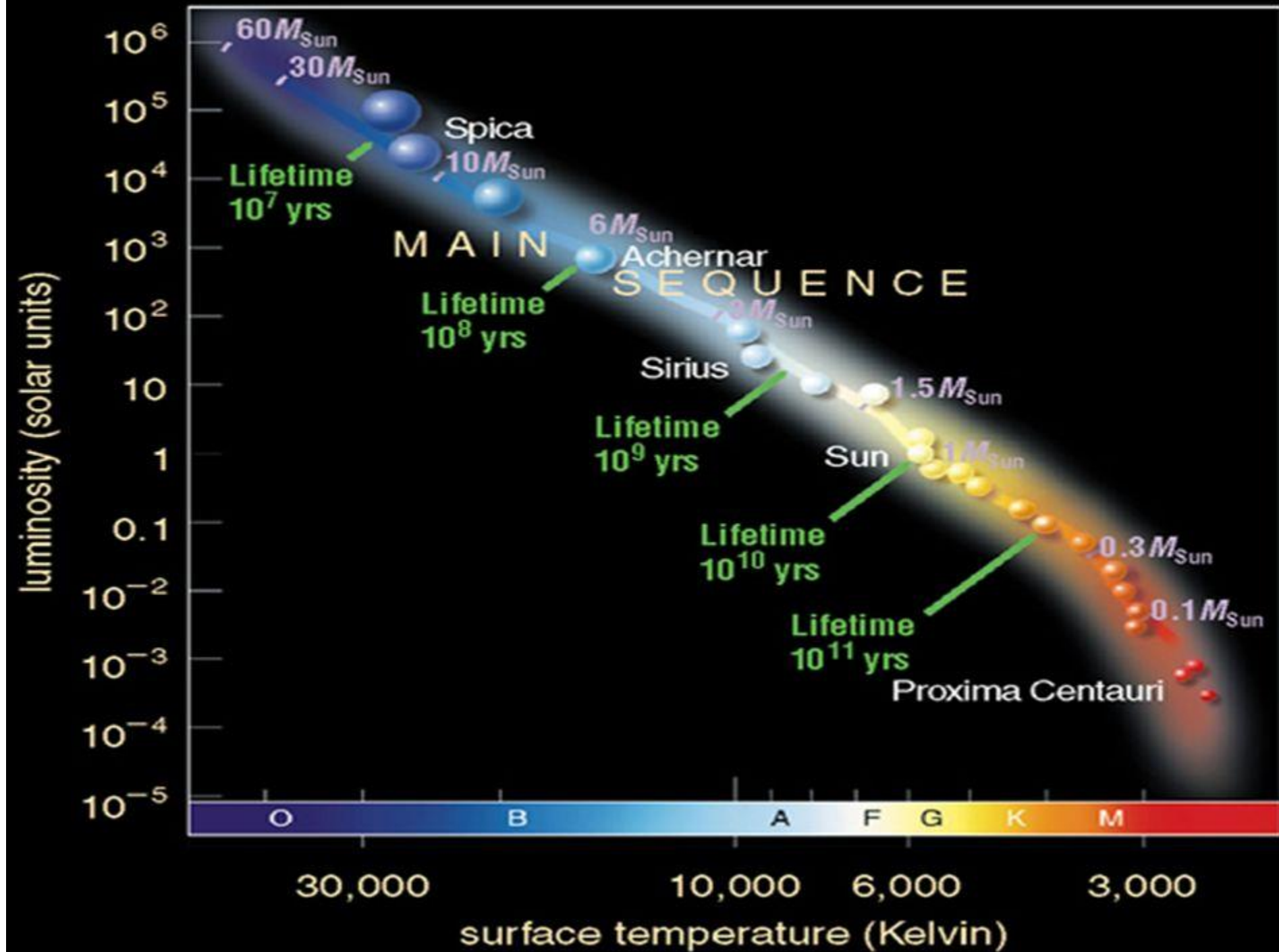
Mercury



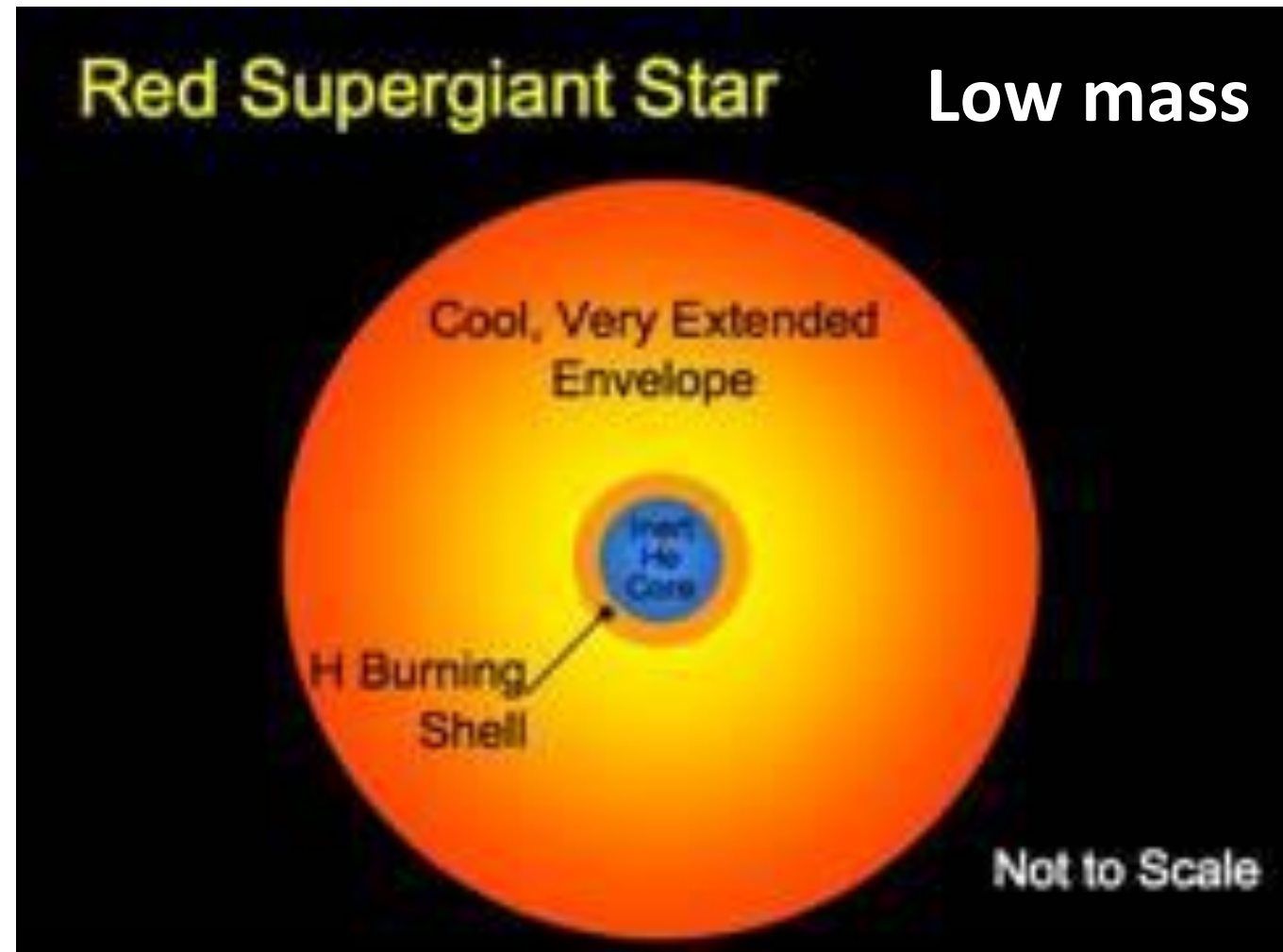
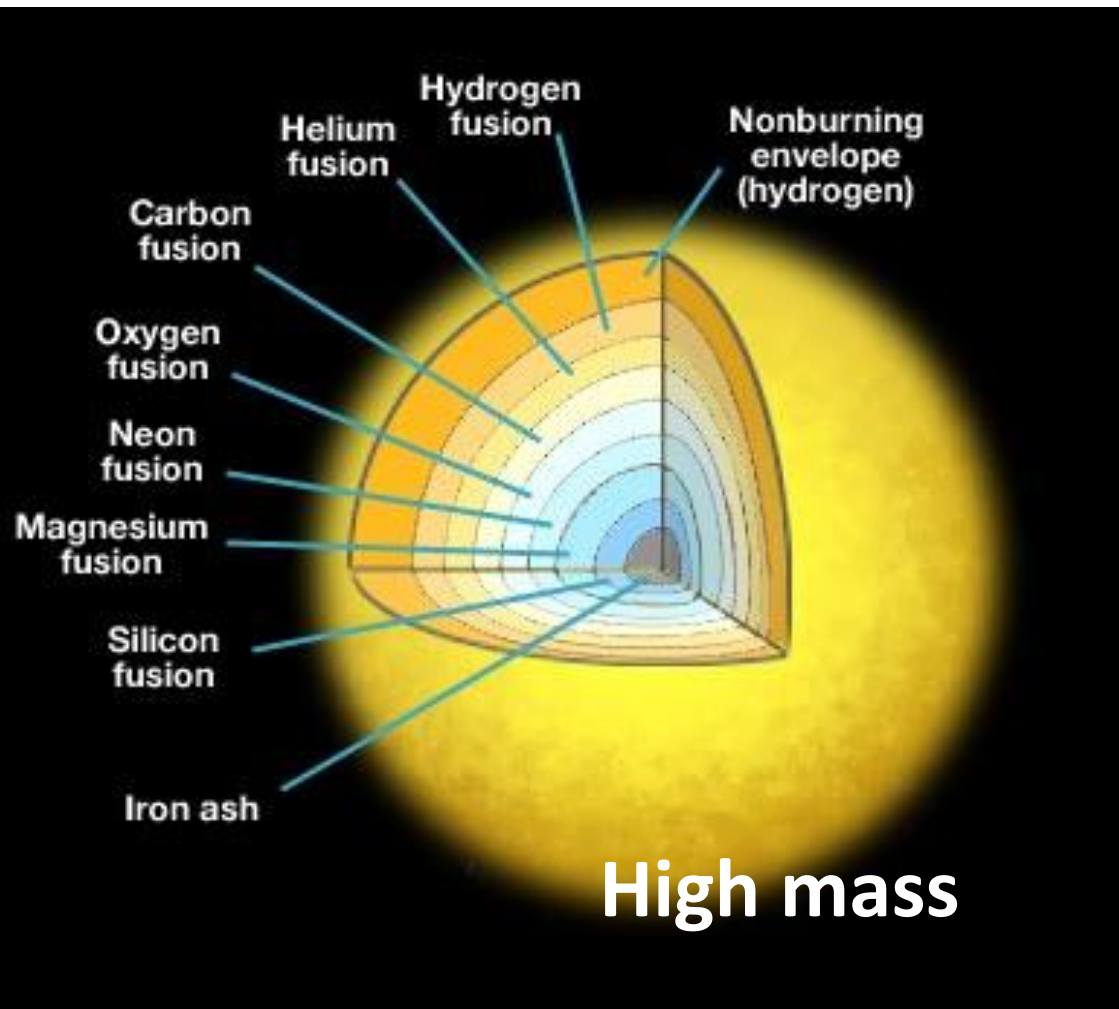
Lifetimes of stars

Big stars have much shorter lives

Why?

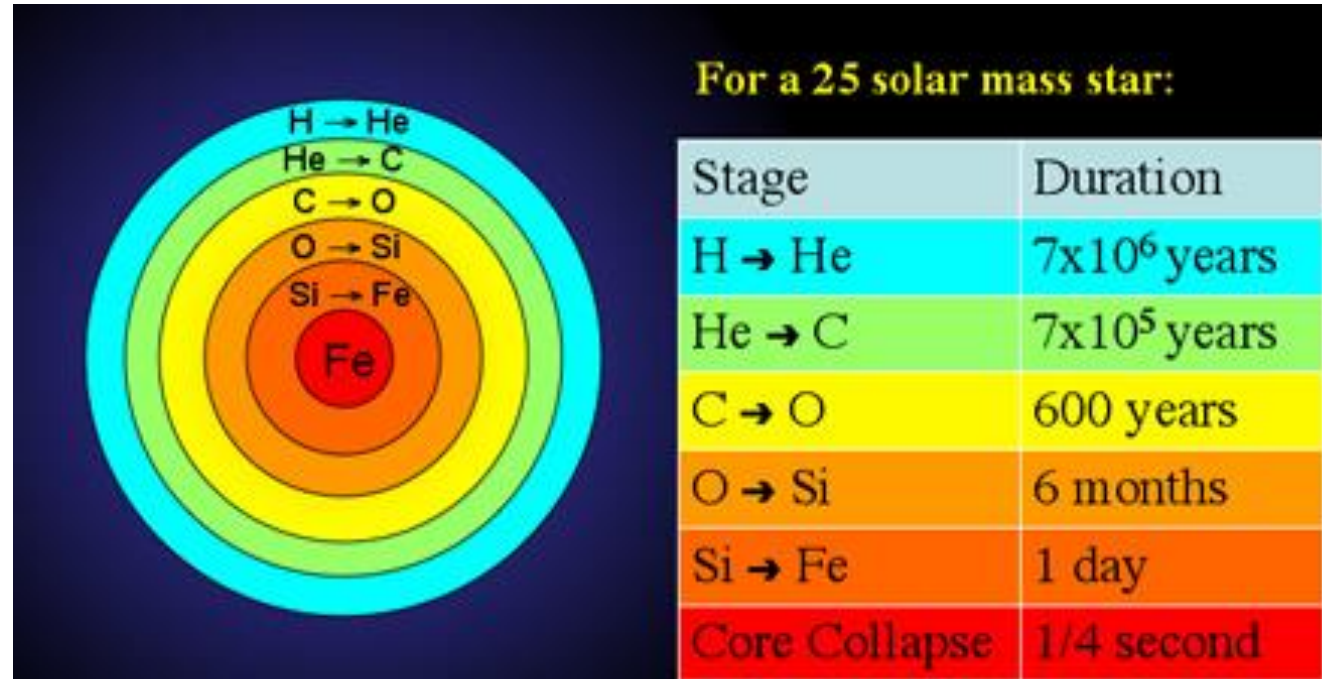
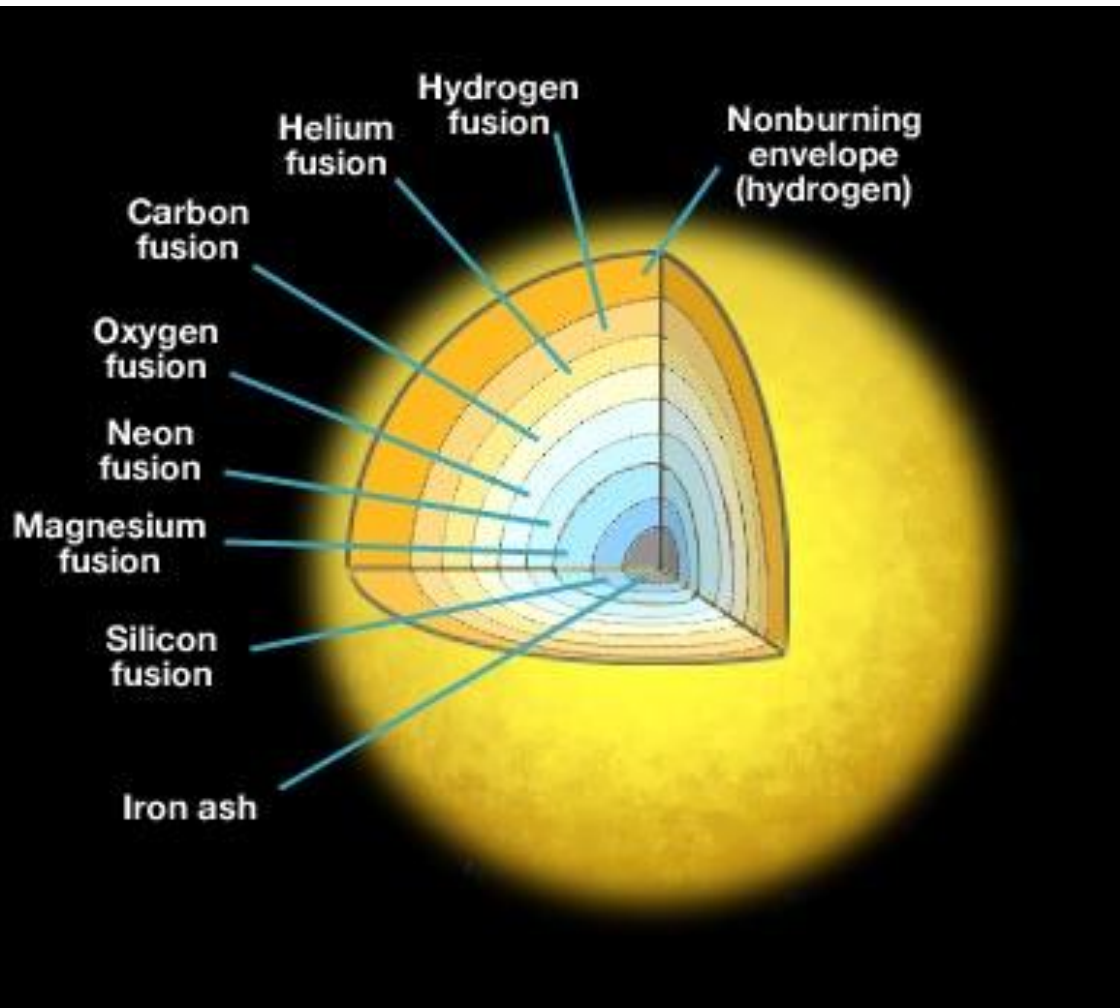


High-mass stars use higher mass elements as nuclear fuel



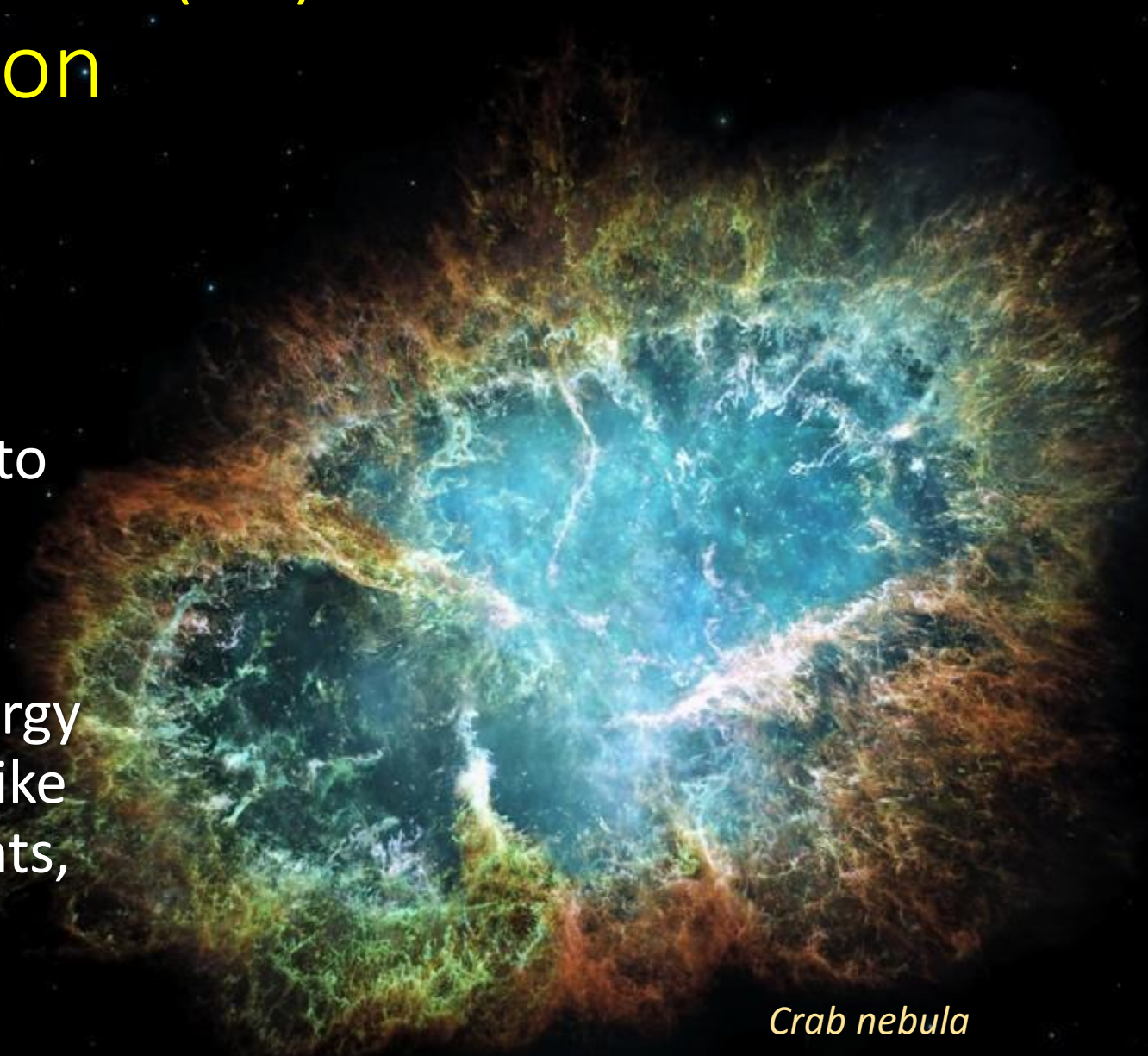
Fusion in high-mass stars

- Go through burning elements in series of steps to make new cores over time



Rapid collapse of iron (Fe) core leads to supernova explosion

- Expands at 10000 km/s
- Releases all the elements into space
- Explosion has such high energy it can fuse heavy elements like Fe into even heavier elements, like Au, Pb all the way to U

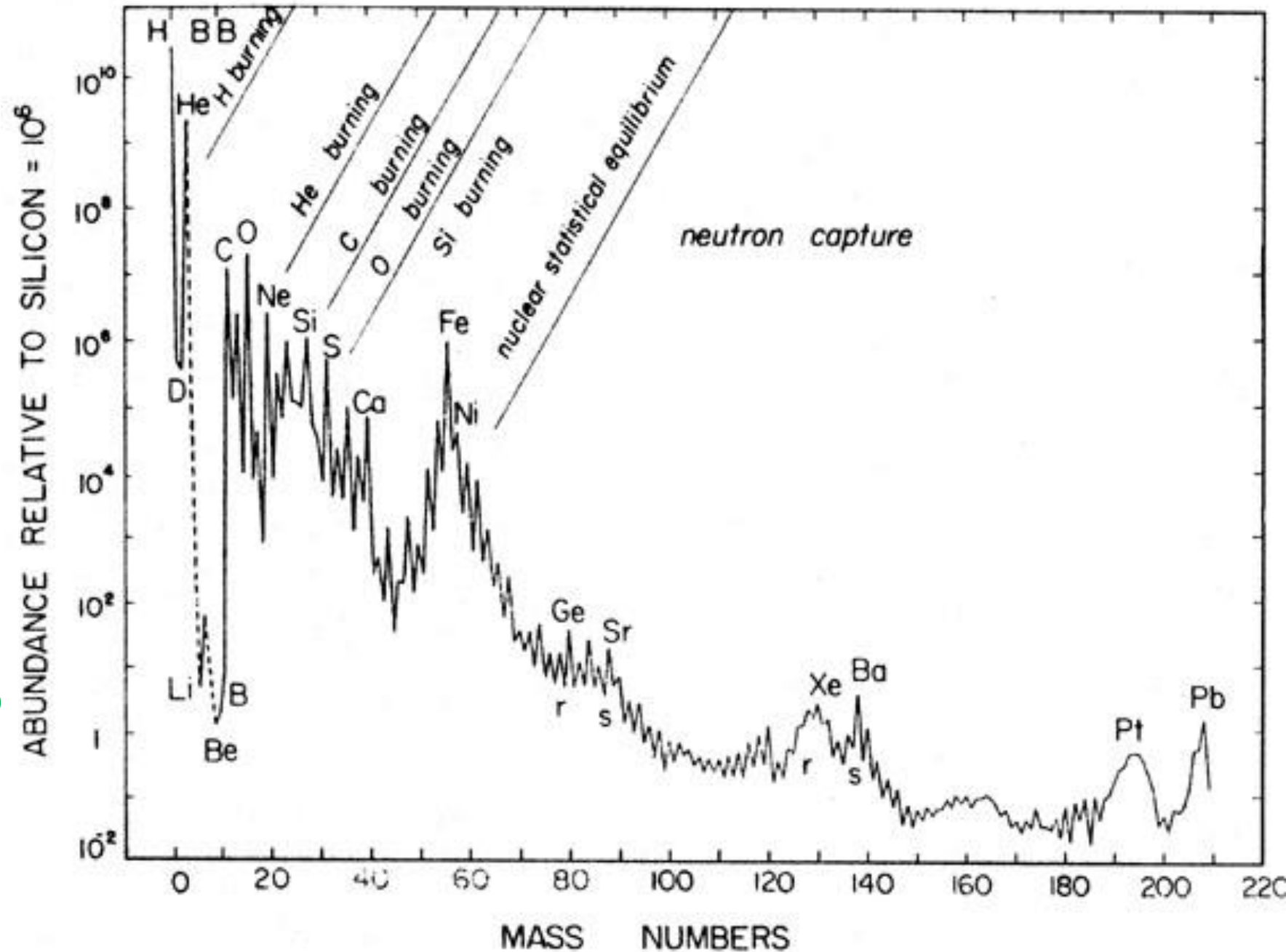


Crab nebula

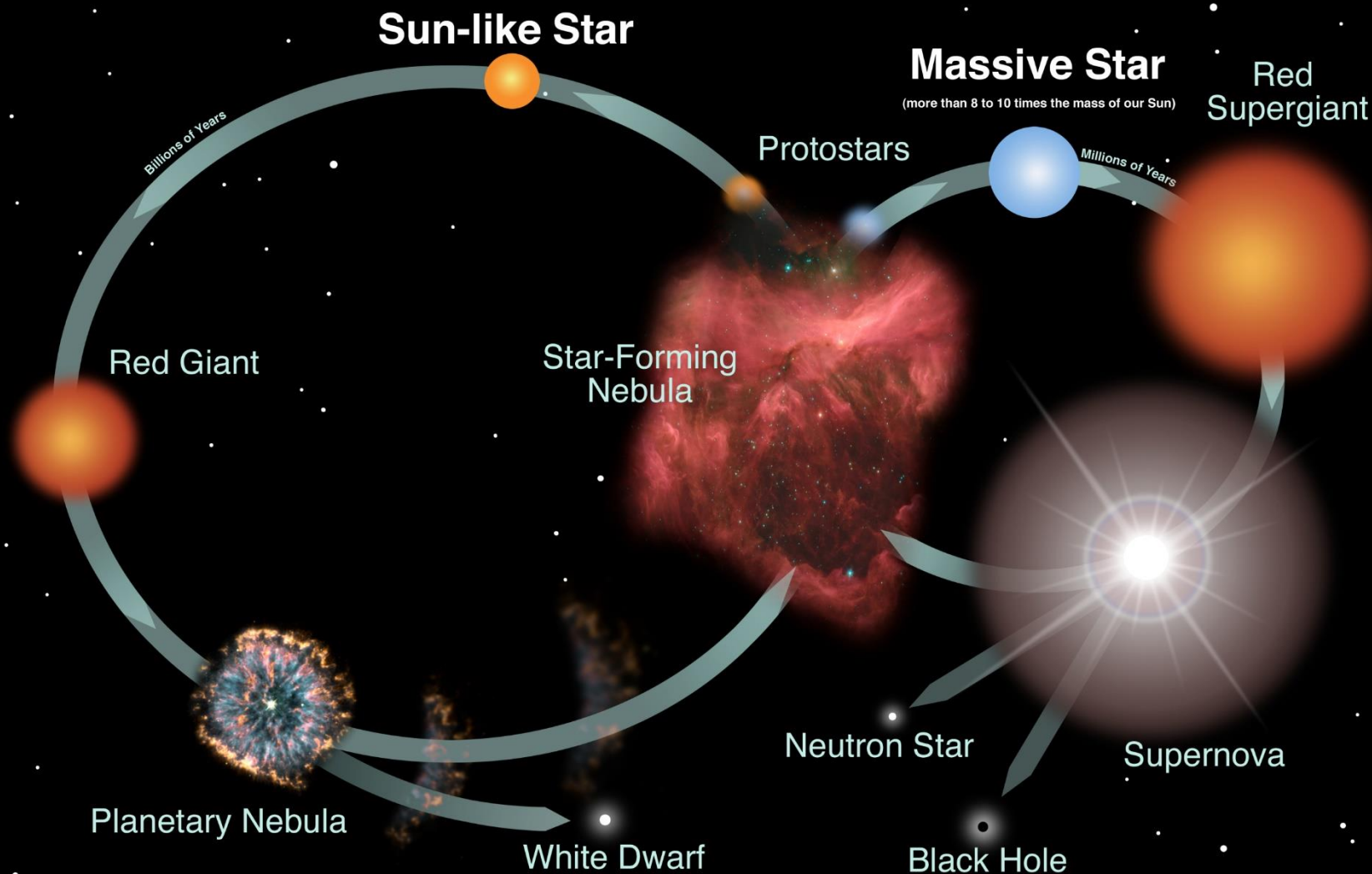
Thank you
nuclear fusion



All the elements we
know of and are made
of were manufactured
in stars and released to
space in supernova
events



Comparing lives of low and high mass stars



the lives of stars